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than the longer one. This shows that, in these two telescopes, the photographic brightness of the stars is proportional to

$$\sqrt[4]{f}$$
 or to  $\sqrt[3]{f}$ .

The disc of a given star was larger in the longer telescope about in the proportion of 50:45.

This proportional size agrees with the result of observation when we assume that like quantities of light are spread over discs of different size—namely, that the shorter telescope gives stars  $\frac{2}{10}$  of a magnitude fainter than the other. Professor Dunér therefore concludes that the different brightness in the two instruments depends only on the fact that the maker did not succeed in getting the star-images so small in the telescope of longer focal length and has nothing to do with the relative foci, as such.

E. S. H.

## ELEVATIONS AT MOUNT HAMILTON.

We are often asked for the exact elevations of points on Mount Hamilton. The following data are derived from a survey made by students of the University of California in 1887, under the direction of W. G. RAYMOND, then their Instructor in Engineering.

•	Feet above Sea.
Top of the wooden cover of the reservoir Copernicus	4383.89
Top of the wooden cover of the reservoir Kepler	4256.28
Highest point of the peak Tycho Brahe	4214.76
Marble floor of the Lick Observatory	4209.46
Top of masonry of the reservoir Huyghens	4178.43
Top of masonry of the reservoir Aquarius	3843.99
Oak tree at Joaquin Springs (B. M.)	
Smith Creek (B. M.)	2146.2
Hall's Valley (B. M.)	
Summit between Hall's Valley and Grand View (B. M.).	
Grand View House (B. M.)	1500.5
Junction House (B. M.)	
San José; S. P. R. R. track at station (assumed)	88.7
	E. S. H.

## METEOR FALL IN ARIZONA.

"Tucson, Sept. 8, 1891.—A meteor of unusual brilliancy and size passed over Tucson at 12:25 last night. It was first seen in the southeast with a long tail of bluish-green fire. Afterward, when seen north of the city, it seemed reduced in size. The light was as brilliant as calcium when seen close to the ground,